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1 Eddy covariance flux measurements of gaseous
2 elemental mercury using cavity ring-down
3 spectroscopy

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6

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12 Keywords: Eddy covariance flux measurements, gaseous elemental mercury (GEM),
13 pulsed cavity ring-down spectroscopy (CRDS), modified Bowen ratio, dynamic flux
14 chamber

15 Abstract

16 A pulsed cavity ring-down spectroscopy (CRDS) system for measuring atmospheric
17 gaseous elemental mercury (GEM) concentrations at high temporal resolution was used
18 to successfully conduct eddy covariance (EC) flux measurements of GEM. GEM fluxes
19 from soils are important to understand because they can originate from natural sources,
20 legacy emissions, and re-emission of previously deposited anthropogenic pollution,
21 which together exceed primary anthropogenic emissions. Eddy covariance flux
22 measurements require sub-second concentration measurements, hampering the
23 measurement of GEM fluxes due to slow response instrumentation. Measurements took
24 place near Reno, Nevada in September and October 2012 encompassing natural, low-
25 mercury (Hg) background soils and Hg-enriched soils. During nine days of measurements
26 with deployment of Hg-enriched soil in boxes within 60 m upwind of the EC tower, clear
27 covariance of GEM concentration and vertical wind speed was measured, showing that
28 EC fluxes over an Hg-enriched area were detectable. During three separate days of flux
29 measurements over background soils (without Hg-enriched soils), no covariance was
30 detected – indicating fluxes below the detection limit. When fluxes were measurable,
31 they strongly correlated with wind direction; and highest fluxes occurred when winds
32 originated from the Hg-enriched area. We also measured fluxes with the modified Bowen
33 ratio (MBR) and a dynamic flux chamber (DFC). Comparisons between the three
34 methods generally showed good agreement in direction (e.g., emission or deposition) and
35 magnitude of fluxes measured, in particular when measured fluxes originated from within
36 the area of Hg-enriched soils. This study demonstrated that a CRDS system can be used

37 to measure GEM fluxes over Hg-enriched areas, with a conservative detection limit
38 estimate of $32 \text{ ng m}^{-2} \text{ hr}^{-1}$, equivalent to the lowest measured absolute flux.

39

40

41 Introduction

42 Long-range atmospheric transport and deposition contributes to significant
43 environmental mercury (Hg) loading even in remote ecosystems¹⁻³. Large uncertainties
44 remain in our understanding of deposition processes, in particular of gas-phase dry
45 deposition which contributes significantly to deposition loads⁴. Gaseous elemental
46 mercury (GEM), the dominant atmospheric form of Hg, is semi-volatile and subject to
47 atmospheric deposition and re-emission leading to a complex pattern of net Hg surface-
48 atmosphere exchange⁴⁻⁶. Processes contributing to GEM re-emission include
49 photochemical reduction⁷⁻¹⁰ as well as other abiotic and biotic processes¹¹⁻¹⁵. Atmospheric
50 GEM emissions from natural sources, legacy emissions, and re-emission of previously
51 deposited anthropogenic pollution have been estimated to exceed primary anthropogenic
52 emissions by a factor of four¹⁶. GEM is the dominant form of Hg in the atmosphere and
53 is therefore the main pathway for Hg to enter ecosystems where it can then be methylated
54 (MeHg) and become bioavailable. Once Hg becomes bioavailable it can bioaccumulate
55 up the food web where it can reach concentrations in marine and terrestrial species
56 known to cause health problem in wildlife and humans¹⁷⁻¹⁹.

57 Several methods for measuring net GEM exchange between surfaces and the
58 atmosphere have been employed, many of them have a number of shortcomings. The
59 most widely used technique, the dynamic flux chamber (DFC), is an easy-to-deploy and
60 relatively inexpensive method that is used to measure the exchange of GEM over a small
61 area (usually $<1 \text{ m}^2$). The DFC method is, however, highly sensitive to pressure
62 differences and flow rates through the chamber²⁰; and placement of the chamber also
63 impacts ambient conditions such as temperature, radiation, humidity, and turbulence

64 levels²⁰⁻²⁴. Micrometeorological techniques, including the modified Bowen ratio (MBR),
65 aerodynamic and relaxed eddy accumulation methods, also have been used to measure
66 GEM fluxes^{21, 25-31}. These methods can be used to measure surface fluxes in situ without
67 modification of environmental conditions, but have other limitations; in particular, the
68 MBR and aerodynamic methods assume equal transport characteristics for GEM and a
69 reference scalar (heat, water, or CO₂) to calculate fluxes (e.g., similarity theory) – which
70 is not always valid³². Another limitation is that concentration measurements are required
71 at two heights which may represent different source areas³³ or cause a bias due to the
72 time lag of alternating measurements between the two heights²⁶. The relaxed eddy
73 accumulation method is a direct flux method, long sampling times are needed to acquire
74 enough sample volume to distinguish concentration differences between updrafts and
75 downdrafts^{25, 34, 35}.

76 Many of these issues could be overcome by using the eddy covariance (EC) technique
77 – the only known direct, in situ method of determining fluxes³⁶⁻³⁸ and the current
78 standard for measuring atmospheric fluxes of many other trace gases (CO₂, methane,
79 ozone, and water vapor)^{27, 36, 39}. The technique uses a sonic anemometer to measure high-
80 frequency eddies (small to large changes in vertical wind direction) and a suite of other
81 instruments to measure scalar atmospheric data to determine vertical turbulent fluxes and
82 therefore exchange rates of trace gases over a variety of landscapes and is useful for
83 assessing exchange over a whole ecosystem for hours to years⁴⁰. This is done by
84 analyzing the covariance between the change in vertical wind speed from the mean (zero)
85 value and the change in a gas concentration (GEM) from its mean value^{40, 41}. This
86 analysis allows for detailed spectral analyses of conditions and allows the operator to

87 qualify data as useable based on conditions favorable for micrometeorological
88 measurements, which is not possible for other techniques discussed previously.
89 Limitations of the EC method include the necessity for flat terrain, steady conditions but
90 turbulent winds, and consistent vegetation⁴⁰. EC measurements also require a fast-
91 response (sub-second) field-deployable sensor, which until recently has not been
92 available for GEM. To test the application of EC for GEM flux measurements, we
93 deployed a newly developed pulsed cavity ring-down spectroscopy (CRDS) system^{42, 43}
94 capable of a measurement resolution of 25 Hz for GEM surface flux measurements with
95 the EC technique. EC flux measurements were conducted over natural low-Hg
96 background soils and over artificially Hg-enriched soil and compared to GEM fluxes
97 measured with the MBR and DFC methods. The CRDS system was used to measure
98 GEM concentrations at 25 Hz. These measurements were then averaged over half hour
99 periods to calculate average GEM fluxes.

100

101 2. Methods

102 2.1 Field site

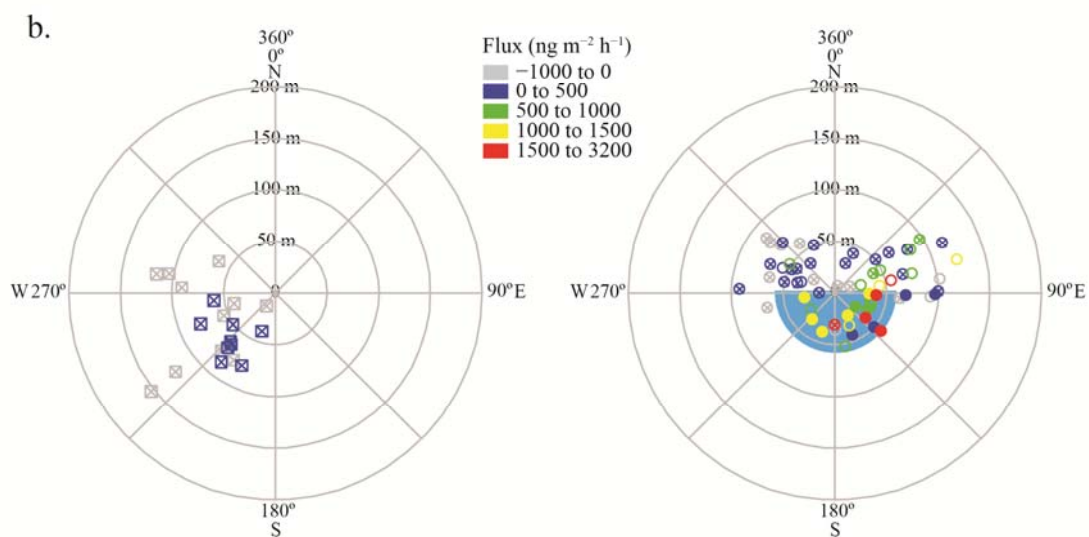
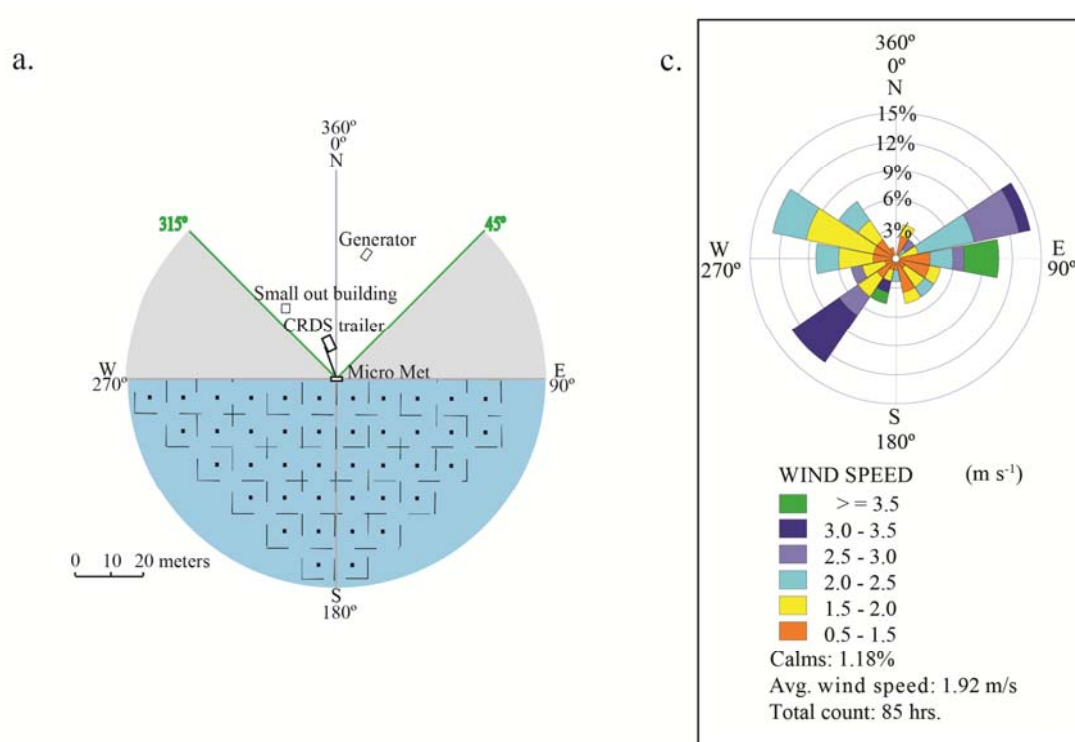
103 Measurements were performed from September 14 to October 1, 2012 at a site 32 km
104 northeast of Reno, Nevada, USA (39° 41' N, 119° 32' W; elevation 1570 m) over a flat
105 area (0.61 km²) of mixed bare soil and shrubs in the Northern Desert Shrub Zone.
106 Surrounding vegetation consisted mainly of sagebrush (*Artemisia tridentata*), rabbitbrush
107 (*Ericameria nauseosa*), and juniper species (*Juniperus spp.*), (Supp. Info. S1). The
108 micrometeorological flux tower was located in the northwest corner of a 50 x 90 m² flat,

109 non-vegetated area with a small hill (elevation 35 m) starting 300 m to the southwest of
110 the tower (Supp. Fig. S1a).

111 GEM flux measurements were first made at the site over soils with low background Hg
112 concentrations (average soil Hg concentration: $0.11 \pm 0.04 \mu\text{g g}^{-1}$) from September 14 to
113 18, termed “background” measurements. From September 19 to September 29, plastic
114 storage boxes containing Hg-enriched soil were deployed as additional GEM sources
115 upwind, in the footprint area (cf. Section 2.4) of the micrometeorological towers, termed
116 “Hg-enriched” measurements (Fig. 1a and Supp. Fig. S1). For this, 41 plastic storage
117 boxes (58 cm x 41 cm) were deployed filled with a layer of Hg-enriched soil (average Hg
118 concentration: $84.5 \pm 5.25 \mu\text{g g}^{-1}$) in a semi-circle around the micrometeorological towers
119 (between 90° and 270° based on predominant wind direction in the area) equally spaced
120 with 10 m between each box (Fig. 1a; Supp. Info. S.1). From September 29 to October 1,
121 the boxes filled with Hg-enriched soil were removed for renewed background
122 measurements (only atmospheric concentrations were measured, no fluxes). The sonic
123 anemometer of the EC system was positioned, due to tower orientation, toward the south-
124 southeast at an angle of 172° allowing unobstructed wind direction and ideal
125 measurements from 82° to 262° , similar to the deployed Hg-enriched soil boxes (see Fig.
126 1a). All flux measurements were made between 12:00 and 18:00 Pacific Daylight Time,
127 the time period when winds typically originated from the $90\text{-}270^\circ$ sector (see Supp. Info.
128 S.2, and Supp. Fig. S.2).

129

130 **Figure 1**



131

132

133 2.2 Flux measurements

134 2.2.1 *Eddy covariance*

135 Three measurement towers were deployed 10 m southeast of a trailer that contained the
136 CRDS system^{42, 43} and auxiliary analyzers. One tower supported a three-dimensional
137 (3D) ultrasonic anemometer for measurement of the three wind components and the
138 speed of sound (CSAT3, Campbell Scientific, Logan, UT, USA) mounted 75 cm above
139 the ground along with an open-path infrared gas analyzer (IRGA, LI-7500, LiCor Inc.
140 Lincoln, NE, USA) for CO₂ and water vapor concentration measurements (Supp. Info.
141 S.3). The open-path IRGA and the air sampling inlet to measure GEM and ozone were
142 mounted 20 cm below the center of the sensing volume of the sonic anemometer, with the
143 IRGA positioned 10 cm east and 20 cm north of the air inlet. The air inlet was connected
144 by 20 m Teflon® tubing (4.76 mm ID) to the mobile measurement trailer (Supp. Fig.
145 S.1b). In addition, the sample air was routed through a set of glassware and ovens for
146 ozone decomposition prior to measurements⁴² and to the closed-path, 1 m quartz-coated
147 stainless-steel measurement cavity of the CRDS system at a flow rate of 8 L min⁻¹ for
148 laminar flow (Reynolds number of ~1800). After the CRDS cavity, flow was distributed
149 to an Hg vapor analyzer (Model 2537B, Tekran Inc., Toronto, Canada) which was used to
150 correct CRDS GEM concentrations at 2.5 min time intervals, a closed-path IRGA (LI-
151 7000, LiCor Inc.; Supp. Info. S.3), and an ozone analyzer (Model 205, 2B Technologies).
152 The trailer also housed two additional Model 2537B analyzers to measure GEM
153 concentrations for the MBR and DFC methods (see Supp. Info S.4).

154

155 2.2.2 *Modified Bowed Ratio*

156 In order to compare EC fluxes with other techniques, we also measured GEM fluxes
157 with the MBR method. Since we were measuring in an arid environment, we chose to use
158 sensible heat as the reference scalar for this technique. A third tower was used for the
159 gradient setup where GEM concentrations and temperature were monitored at 0.58 and
160 1.09 m above the ground. Temperature at each height was continuously measured (1 min
161 average) with two 1000 Ohm platinum resistance thermometers (PRTs; $\pm 0.01^\circ \text{C}$;
162 Thermometrics Corporation; Los Angeles, CA, USA). Each set of PRTs was placed in a
163 fan-aspirated radiation shield (Model 076B; Met One Instruments; Grants Pass, OR,
164 USA). Both heights were sampled for GEM by a single Tekran 2537B, located inside the
165 adjacent mobile trailer and connected by Teflon® tubing, over five minute intervals (two
166 2.5 min samples on each trap consecutively). A Tekran automated Dual Switching unit
167 (Model 1110; Tekran Inc.; Toronto, Canada) was used to switch between the two sample
168 inlets. We used a pump to flush the inlet that was not being sampled to avoid stagnant air
169 in the sample lines between sampling. A data logger (CR3000; Campbell Scientific;
170 Logan, UT, USA) was used to collect data from the PRTs. More information can be
171 found in Supp. Info. S.4.

172

173 *2.2.3 Dynamic flux chamber*

174 Further, a dynamic flux chamber constructed of Teflon was used to quantify GEM
175 fluxes over a surface area of 0.036 m^2 . The flux chamber – provided by the laboratory of
176 M. Gustin at the University of Nevada, Reno – had a height of 6.5 cm, 2 L of volume,
177 and 1 cm diameter holes every 2.5 cm around the chamber circumference located 2 cm
178 above the surface to allow unrestricted airflow at 1.2 L min^{-1} during background

179 measurements and 4 L min^{-1} during measurements over a source tub²⁰. The chamber was
180 positioned over an area of open soil with minimal disturbance. GEM concentrations
181 inside and outside the chambers were measured by a second 2537B Tekran Hg analyzer
182 connected by 25 m of Teflon tubing. Switching between chamber inlet and outlet air was
183 controlled with a Tekran Model 1115 multi-port switching unit. More information can be
184 found in Supp. Info. S.4.

185

186 *2.3. EC flux calculations*

187 We calculated half-hour average fluxes for GEM as well as sensible and latent heat
188 based on the covariance between turbulent fluctuations of the vertical wind speed and the
189 scalar mixing ratios using Reynolds (block) averaging^{27, 37, 38, 44, 45}. Negative fluxes
190 reported here represent transport from the atmosphere toward the surface (deposition),
191 while positive fluxes represent emission. For details on the post-processing of EC fluxes
192 see Supp. Info. S.3. We assigned half-hour fluxes quality control (QC) flags: 1 (good), 2
193 (usable), and 3 (not usable). The following criteria were used to assign QC flags: (i)
194 following the procedure outlined by Ruuskanen³⁶, the GEM cross-correlation with
195 vertical wind speed (covariance of vertical wind speed and GEM concentrations with the
196 latter shifted in time sample-by-sample) was visually assessed and given QC flag 1 when
197 a well-defined peak in the cross-correlation was found within the searched time window
198 (± 10 s); QC flag 2 was assigned when a less defined, but still detectable, peak was found;
199 and QC flag 3 was assigned when no peak was observed; (ii) optimal wind directions for
200 the sonic anemometer setup according to the manufacturer's specifications (QC flag 1: $>$
201 90° and $< 270^\circ$; QC flag 2: $> 45^\circ$ and $\leq 90^\circ$ plus $\geq 270^\circ$ and $< 315^\circ$; QC flag 3: all other

202 cases); (iii) following ⁴⁶⁴⁵⁴⁴⁴³⁴³⁴³⁴²⁴¹⁴¹McMillen⁴⁶, QC flag 3 was assigned when the
203 third rotation angle exceeded $\pm 10^\circ$ (QC flag 1 in all other cases); (iv and v) following
204 Mauder et al.⁴⁷, QC flag 1 was assigned when the deviation of the stationarity test (a test
205 of how variable the covariance is over time)⁴⁷, for the GEM flux or the integral similarity
206 characteristics (a comparison with an accepted model of turbulence) was $<30\%$, QC flag
207 2 was assigned for a deviation $\geq 30\%$ and $<100\%$, and QC flag 3 was assigned in all other
208 cases. For each half-hour period, the overall QC flag was conservatively derived as the
209 maximum QC flag of the five criteria above (Supl. Fig. S.3).

210

211 *2.4 Footprint analysis*

212 We estimated the EC flux source area (footprint) for each half-hour period with the
213 Hsieh et al.⁴⁸ footprint model (Supp. Info. S.5, Fig. 1.b). The flux footprint is an estimate
214 of the actual source area contributing to the GEM flux based on atmospheric conditions
215 (e.g., wind direction, wind speed, stability), surface properties (e.g., roughness), and
216 height of the measurement instruments^{30, 31, 33, 48}. This was important for estimating the
217 source area contributing to flux measurements as Hg-enriched surface sources were
218 placed within 60 m of the tower and within an angle of 90° to 270° . The flux footprint is
219 expressed as $X_{90\%}$ values representing the upwind distance from the tower within which
220 90% of the flux originated (source area). In addition, the concentration footprint for each
221 of the two intake heights of the MBR system was estimated using a 3-D backward
222 Lagrangian footprint model (LPDM-B)³³. In contrast to the flux footprint, which was
223 calculated for each available half-hourly period, the LPDM-B model was run with typical
224 atmospheric conditions during the entire measurement campaign (i.e., mixing-layer

225 heights of 3500 and 5000 m, Monin-Obukhov lengths of -5 and -15 m, and convective
226 velocity scales of 2 and 3 m s⁻¹, respectively). The roughness length was set to 0.01 m for
227 bare soil in the LPDM-B model, which corresponds to the average value calculated from
228 the logarithmic wind law under near-neutral atmospheric conditions when winds
229 originated from the 90-270° sector.

230

231 3. Results and discussion

232 3.1. High-frequency signal analysis

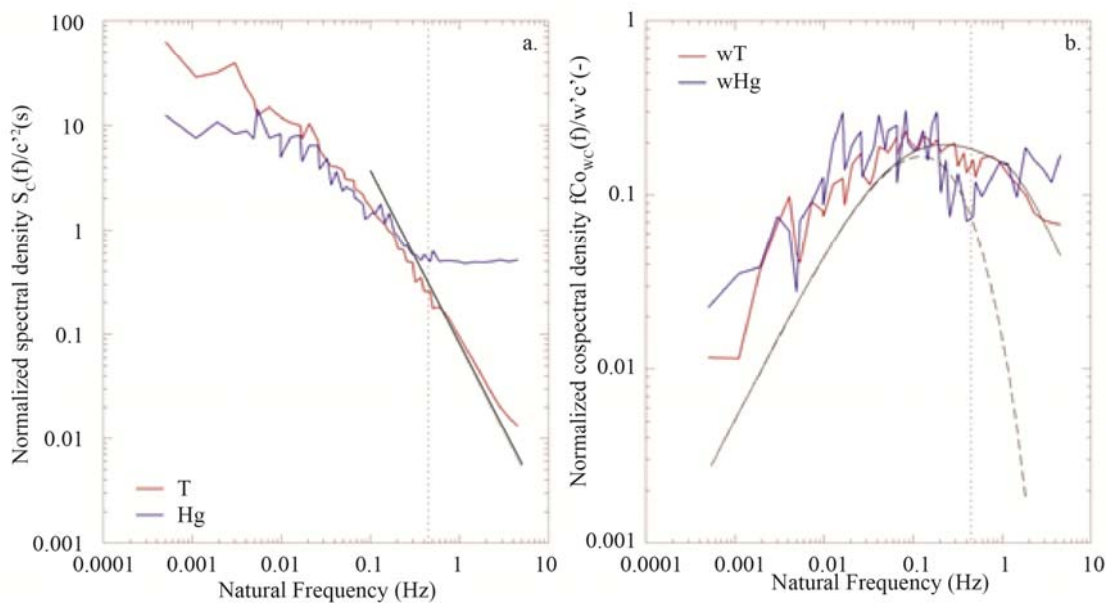
233 We thoroughly checked collected GEM concentrations for data quality before
234 calculation of EC fluxes, including Allan variance plots (Supp. Fig. S.4), (co-)spectral
235 (Fig. 2) and cross-correlation analysis (Fig. 3); these served to validate GEM
236 concentration measurements and supported processing data prior to calculating fluxes (as
237 detailed below). We used Allan variance plots to assess the stability of GEM
238 concentration measurements through time. An example, shown in Supp. Fig. S.4,
239 demonstrates that the variance of high-resolution GEM measurements (red line) during
240 stable background conditions decreased with a slope inversely proportional to the
241 integration time until approximately 50 s. The Allan variance started to increase after an
242 integration time of approximately 200 s with a slope proportional to the integration time.
243 The latter is an indication of linear drift of the CRDS system causing an increase in the
244 GEM variance at longer integration times⁴⁹. In order to account for this instrument drift,
245 which would result in an overestimation of GEM flux calculations, we corrected CRDS
246 GEM signals using GEM concentrations measured downstream of the CRDS with a
247 slower-time response (2.5 min) Model 2537B Tekran Hg analyzer. To this end, we

248 calibrated CRDS GEM data every 2.5 min with concentrations measured by the Model
249 2537B. As shown in Supp. Fig. S.4, this procedure (blue line) effectively removed low-
250 frequency drifts, causing the variance to decrease with a slope inversely proportional to
251 the integration time through 1000 s.

252 We used spectral and co-spectral analyses to compare the quality of the CRDS GEM
253 signals with the temperature data as measured by the sonic anemometer in the frequency
254 domain⁵⁰. Power spectra and co-spectra, typical for QC 1, are shown in Fig. 2. This
255 comparison between temperature and GEM power spectra showed that the GEM signal
256 started to be dominated by noise above frequencies of 0.45 Hz, causing it to deviate from
257 the expected $f^{5/3}$ (where f is frequency) decay in the inertial sub-range (solid line Fig. 2a)
258 and the temperature spectrum measured with the sonic anemometer⁵¹. The co-spectra of
259 vertical wind speed with temperature and GEM, respectively, generally agreed in the
260 lower frequency range and up to around 0.2 Hz, indicating that the removal of the linear
261 drift in the GEM signals was successful (Fig. 2b). At frequencies around 0.2 Hz, the
262 GEM co-spectrum started to roll off, which is indicative of low-pass filtering by the
263 measurement system (in particular attenuation of concentration fluctuations down the
264 sampling tube)⁵². The dashed line in Fig. 2b represents the co-spectral reference model
265 convolved with a series of transfer functions which account for all sources of low-pass
266 filtering (Supp. Inf. S.3). Our approach of correcting for low-pass filtering was able to
267 reproduce (and thus allowed correcting for) this flux loss^{51, 52}. At frequencies >0.45 Hz
268 (which corresponds to ca. 12% of total flux), the noise observed in the GEM spectra (Fig.
269 2a) compensated for the low-pass filtering and even caused an increase in co-spectral
270 density, while the temperature co-spectrum further decreased in the inertial sub-range as

271 expected⁵¹ (Fig. 2b). In order to remove the noise in the high-frequency part of the GEM
 272 co-spectrum, we used a low-pass Finite Impulse Response (FIR) filter⁵¹ with a time
 273 constant of 2 s to filter out any unwanted contributions at frequencies $>0.45\text{Hz}$. The
 274 missing high-frequency ($>0.2\text{ Hz}$) flux contribution due to low-pass filtering by the
 275 measurement system and the low-pass FIR filter ($>0.45\text{ Hz}$) was then back-corrected
 276 based on the reference model co-spectrum of Kaimal and Finnigan⁵¹. The resulting
 277 average frequency response correction factor was 1.89 (range 1.29–2.58).

278 **Figure 2**



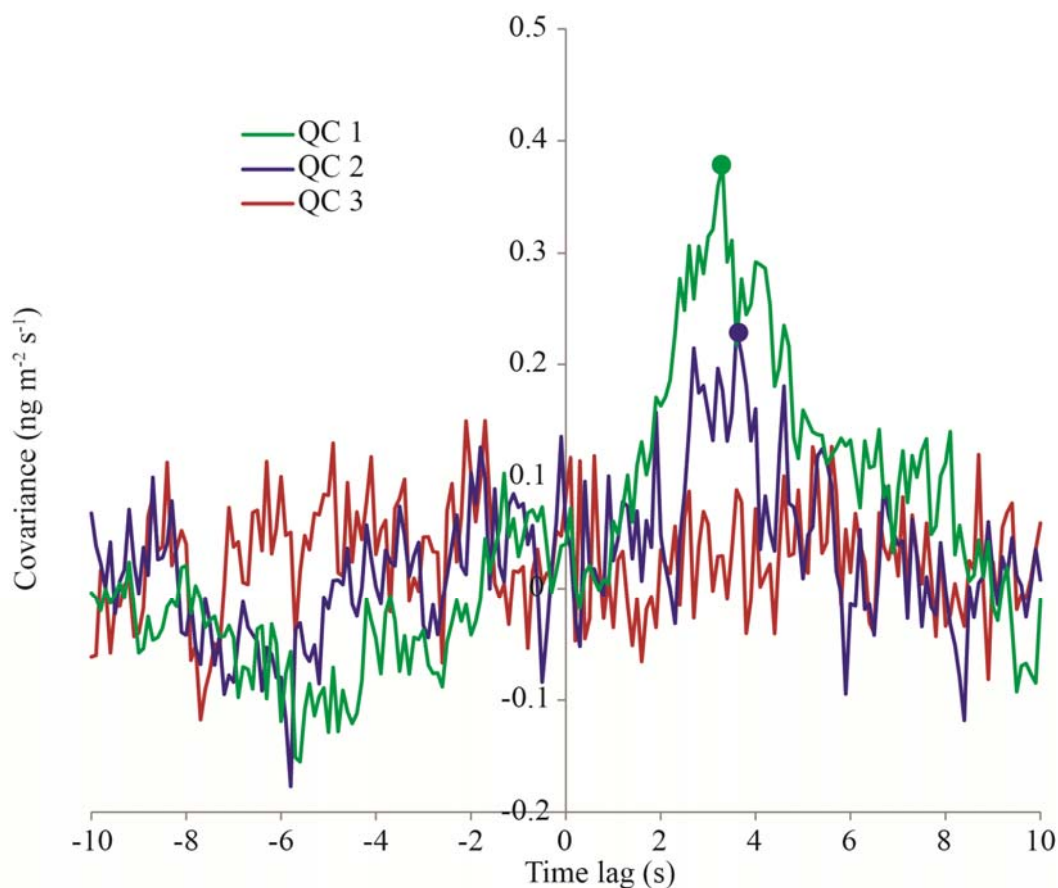
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280

281 We used a cross-correlation analysis to identify whether a covariance existed between
 282 vertical wind speed and GEM concentration for each half hour sampled. Following
 283 Ruuskanen et al.³⁶, a well-defined peak in the cross-correlation analysis (Fig. 3) indicated

284 a significant correlation between vertical wind speed and GEM concentration from which
285 a flux was calculated. Figure 3 demonstrates examples of a well-defined peak in the
286 cross-correlation between GEM and vertical wind speed (assigned QC 1) as well as a
287 lower quality peak (assigned QC 2), which was associated with higher uncertainty. A
288 lack of cross-correlation (assigned QC 3) indicated fluxes close to zero, which may be
289 because no fluxes existed or because fluxes were below the detection limit of the system.
290 The time lag shown in the cross-correlation analysis, as indicated by the shift in the peak
291 covariance, at roughly 3.3 s represents the time for sample air to travel from the sample
292 inlet on the EC tower (where vertical wind speed was measured) to the CRDS system
293 inside the measurement trailer (where GEM was measured). This time lag, which was
294 consistent throughout the measurements, was used to properly align all vertical wind
295 speed and GEM concentration data for covariance calculations.

296 **Figure 3**



297

298 Of a total of 85 half-hour measurements, 17%, 21%, and 62% fell into the overall QC
299 classes 1, 2, and 3, respectively (Supp. Fig. S.3), and 57% of the data in QC 3 was due to
300 lack of a clear cross-correlation. During sampling under background conditions, all
301 available half-hour fluxes failed to meet the cross-correlation criterion (i.e., fluxes close
302 to zero and/or below detection limit). During deployment of the Hg-enriched soils, 43%
303 of the data failed to meet the cross-correlation criterion, which usually was associated
304 with air masses originating from outside the Hg-enriched soil area. Overall, the wind
305 direction, stationarity, integral turbulence, and third rotation angle criteria were not
306 fulfilled in 7, 11, 0, and 1% of all cases.

307 Several methods were used to evaluate the detection limit of the EC flux system. The
308 smallest, absolute detected GEM flux with a clear cross-correlation was $32 \text{ ng m}^{-2} \text{ h}^{-1}$
309 (QC 2), Using the method described in Pihlatie et al. (2005)⁵³, which requires the
310 standard deviation of vertical wind speed and the noise level of the instrument (30 min
311 standard deviation during background measurements), we calculated a flux detection
312 limit of $2\text{-}60 \text{ ng m}^{-2} \text{ h}^{-1}$. Standard deviation of vertical wind speed varied between 0.20-
313 0.45 m s^{-1} . The standard deviation of GEM during background measurements varied
314 between $0.0022\text{-}0.0050 \text{ } \mu\text{g m}^{-3}$. This flux detection limit is similar to the smallest
315 absolute detected GEM flux of $32 \text{ ng m}^{-2} \text{ h}^{-1}$. However, using the power spectra of the
316 background noise (before Hg-enriched soil boxes were in place) and an estimate of the
317 vertical wind power spectra as described in Pattey et al. (2006)⁵⁴, we calculated a flux
318 detection limit of $82\text{-}195 \text{ ng m}^{-2} \text{ h}^{-1}$ for the range of observed wind speeds of $0.3\text{-}4 \text{ m s}^{-1}$.
319 This technique gives a higher flux detection limit. Following Wienhold et al. (2005) a
320 signal-to-noise (S/N) ratio analysis was performed. For the 85 half-hour periods of
321 calculated fluxes, 33% exceed a S/N of 3. In QC 1 S/N varied from 4.1 to 9.5 (100% S/N
322 > 3), QC 2 S/N varied from 2.6 to 9.2 (95% S/N > 3) and QC 3 S/N varied from 1.8 to
323 23.4 (72% S/N > 3). This indicates that for QC 1 and 2 S/N ratios were generally above
324 3. Although some fluxes with acceptable S/N ratios were rejected in QC 3, the use of
325 other quality criteria in this case assures us that these fluxes are not useable (Supp. Fig.
326 S.3 and S.5).

327

328 *3.2. EC GEM fluxes*

329 Mean background afternoon (12:00–18:00) ambient air GEM concentrations measured
330 with a Model 2537B analyzer were 1.42 ng m^{-3} with a range of 1.20 to 2.23 ng m^{-3} prior
331 to the Hg-enriched soil being deployed and 1.53 ng m^{-3} with a range of 1.06 to 2.85 ng m^{-3}
332 after removal. During the three days of background flux measurements, GEM fluxes
333 calculated with the EC method (Table 1; Fig. 1b, left panel) all failed the cross-
334 correlation criteria (QC 3). As a result, all of these fluxes can be considered below the
335 detection limit of the system.

336 When the Hg-enriched soil was deployed (Sept. 19 to 28), mean ambient air GEM
337 concentrations increased to 3.75 ng m^{-3} with a range of 0.68 to 21.40 ng m^{-3} . The LPMD-
338 B footprint model showed that during typical afternoon conditions, the source area ($X_{90\%}$)
339 for 80% of the half-hour EC GEM fluxes originated within 14 to 18 m of the
340 measurement tower, well within the area where Hg-enriched soil sources were deployed
341 (extending 60 m). As expected, the highest fluxes ($1000\text{--}3200 \text{ ng m}^{-2} \text{ h}^{-1}$) commonly
342 were observed during periods when $X_{90\%}$ fell within the Hg-enriched area. As can be seen
343 in Figure 1b, large fluxes ($>1000 \text{ ng m}^{-2} \text{ h}^{-1}$) also were observed on three occasions where
344 $X_{90\%}$ values were beyond, but closely adjacent to, the Hg-enriched soil sources (discussed
345 below). Lower fluxes, below $1000 \text{ ng m}^{-2} \text{ h}^{-1}$ and often below $500 \text{ ng m}^{-2} \text{ h}^{-1}$, occurred
346 mainly when fluxes originated from wind directions outside the Hg-enriched soil sources
347 and during background measurements. GEM fluxes derived with all methods increased
348 (Table 1), and EC GEM fluxes were more consistently classified as QC 1 and QC 2 (Fig.
349 1b, right panel; Supp. Fig. S.3). Figure 1b shows flux magnitudes and the source area of
350 GEM fluxes expressed as $X_{90\%}$ for all 85 half-hour GEM EC measurements.

351

352 **Table 1.** Comparison of fluxes over background and Hg-enriched soils derived using all
 353 three techniques.

Background measurements	Mean	Min	Max	StdError	n	% no cross-correlation
GEM (ng m ⁻³)*	1.48	1.06	2.85	0.01	448	
EC (ng m ⁻² h ⁻¹) **	<MDL***	<MDL	<MDL	<MDL	<MDL	100
MBR (ng m ⁻² h ⁻¹) **	-76	-265	50	17	20	-
DFC (ng m ⁻² h ⁻¹)	0.01	-8.9	6.8	0.09	277	-
Hg-enriched area measurements						
GEM (ng m ⁻³)*	3.75	0.68	21.40	0.09	743	
EC QC 1 (ng m ⁻² h ⁻¹)	1105	71	2339	180	14	0
EC QC 1&2 (ng m ⁻² h ⁻¹)	849	-304	2339	111	32	0
MBR QC 1 (ng m ⁻² h ⁻¹)	1712	-795	3934	367	14	-
MBR QC 1&2 (ng m ⁻² h ⁻¹)	1309	-1026	3934	217	32	-
DFC (ng m ⁻² h ⁻¹)	1105	23	3848	68	150	-

354 *atmospheric concentrations measured by Tekran 2537B

355 **all background EC fluxes were QC 3 and had no cross-correlation, MBR fluxes were below the
 356 minimum resolvable gradient (MRG)

357 ***MDL = 32 ng m⁻² h⁻¹

358

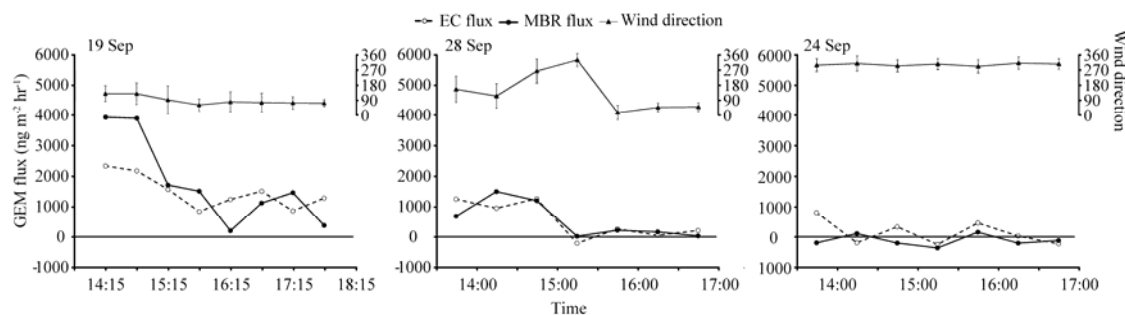
359 In general, GEM fluxes measured with the EC method when the Hg-enriched soil
 360 boxes were deployed had a wide range of values (QC 1&2 standard deviation \pm 626.4;
 361 Supp. Fig S.6). Fluxes previously reported from naturally Hg-enriched soils ²³, fluxes
 362 measured in large ecologically controlled lysimeter laboratories (EcoCELLs) over Hg
 363 amended soils ⁵⁵, and fluxes calculated from DFC measurements using Hg-enriched
 364 substrates ⁵⁶⁻⁵⁸ showed a range that spanned five orders of magnitude in comparison to
 365 our measurements. We believe the measurement set-up contributed most of this
 366 variability (Table 1); and, in particular, the differences between GEM fluxes over the Hg-
 367 enriched soil and background soil area. The Hg-enriched soil boxes were limited to a

368 semi-circle from due east through south to due west (Fig. 1a), and it was likely that
369 variability in EC GEM fluxes was induced by flux source areas switching between
370 background and Hg-enriched areas (e.g., an “edge effect”, during any half-hourly
371 averaging period). Another possibility is that environmental conditions caused some of
372 the variability. This is unlikely since measurements were taken during relatively unstable
373 conditions in afternoon hours, with no major precipitation events occurring.

374 To highlight the edge effect, Fig. 4 shows three days of EC and MBR flux
375 measurements that were impacted differently by the Hg-enriched soil area due to changes
376 in dominant wind directions. On September 19, all measurements corresponded to
377 average wind directions from the Hg-enriched area (between 90° and 270°). Fluxes were
378 highest when the wind direction was centered within the Hg-enriched soil area and as
379 winds moved towards the edge and out of this area, measured fluxes decreased. All
380 measurement points on September 19 were characterized as QC 1 and 2. The edge effect
381 also was obvious on September 28 when dominant winds first originated from the Hg-
382 enriched soil area (first three measurements), but fluxes dropped significantly after the
383 flux source area moved into the background soil area. While the initial three
384 measurement points were characterized as QC 1 and 2, the last four points were
385 characterized as QC 3 due to a lack of a cross-correlation and acceptable wind direction,
386 and were consistent with lower GEM fluxes expected over the background soil area.
387 Finally, on September 24, all measurements were taken when the flux source area was
388 located outside of the Hg-enriched soil area, and fluxes, although relatively small, should
389 be considered with caution during this period since five of the seven points were
390 characterized as QC 3 data.

391 The examples on September 19, 24, and 28 show sensitivity of the EC measurements to
 392 wind directions shifting over the Hg-enriched soil area. It is further possible that
 393 additional flux variability was due to the deployment of discrete Hg sources; these
 394 sources were placed in boxes to avoid contamination of the site with 10 m in between,
 395 and the emissions from these tubs may not have been thoroughly mixed in the
 396 atmosphere prior to measurements at the EC tower, inducing additional variability both in
 397 concentrations and calculated fluxes.

398 **Figure 4**



399

400 3.3 Comparison of GEM flux measurement methods

401 We compared EC fluxes with measurements from two other flux measurement
 402 methods, the MBR method and a DFC. Over background soils, mean GEM fluxes
 403 measured with the DFC (Table 1) were within the range of GEM fluxes measured with
 404 DFCs in other background areas (ranging from -1 to 45 ng m⁻² hr⁻¹)^{24, 59, 60} and over
 405 background soils in Nevada (averaging 1.0 to 2.0 ng m⁻³ h⁻¹; with ranges of -3.7 to 9.3 ng
 406 m⁻² h⁻¹)⁶¹⁻⁶³. The range of measured EC GEM fluxes was substantially higher than those
 407 measured with the DFC and actually showed an average depositional flux across all
 408 background measurements of -43 ng m⁻² hr⁻¹. We considered these measured EC fluxes
 409 over background soils unreasonable (both regarding variability and magnitude) as

410 compared to background soil fluxes reported by other studies and our own background
411 GEM fluxes measured by the DFC method. This is supported by all of the EC fluxes
412 falling into the QC 3 category with no detectible cross-correlation between GEM and
413 wind speed.

414 The issue with many GEM fluxes falling below the detection limit was not unique to
415 the EC method but was also apparent in fluxes measured with the MBR method
416 (averaging $-76 \text{ ng m}^{-2} \text{ hr}^{-1}$). Resolving very small vertical gradients between two
417 measurement heights is challenging over non-contaminated background sites, which was
418 further exacerbated by a short separation (0.51 m) between measurement inlets compared
419 to more commonly used gradients of 2 to 4 m. In this study, the small vertical separation
420 was necessary to restrict flux footprints to the Hg-enriched soil area (60 m upwind).
421 Other GEM fluxes measured with micrometeorological methods over background
422 ecosystems fell in the range of -68 to $34 \text{ ng m}^{-2} \text{ hr}^{-1}$ ^{21, 28, 62, 64} and often showed high
423 variability as well. Calculated minimum resolvable gradients (MRG) between the two
424 GEM inlets in MBR setups are in the range of 0.01 to 0.05 ng m^{-3} in several studies^{21, 31,}
425 ⁶⁴. During our background measurements, MBR gradients averaged 0.04 ng m^{-3} with a
426 range of -0.04 to 0.16 ng m^{-3} , and 50% of the data was below an absolute value of 0.05
427 ng m^{-3} which was the smallest detectable gradient in a previous study²¹.

428 Unlike under background conditions, GEM fluxes measured with the EC, MBR, and
429 DFC methods over the Hg-enriched soil area were in similar ranges, generally in the
430 same direction (emission or deposition), and showed similar average fluxes and temporal
431 patterns (Table 1). Fluxes measured with DFCs have been reported to be lower than those
432 measured with micrometeorological methods^{22, 23}, which was not the case for our

433 measurements. Figure 4 shows a flux comparison between EC and MBR measurements
434 for the three days that were presented above to discuss the effects of wind direction. The
435 patterns show that EC and MBR measurements compared well during these days,
436 showing similar temporal patterns and magnitudes of fluxes. This may indicate that
437 conditions that were favorable or unfavorable for the EC measurements, based on
438 spectral analysis, are similarly favorable or unfavorable for the MBR method. For
439 September 19, when all EC measurements fell into QC 1 and 2, the coefficient of
440 determination, r^2 , of a significant linear regression between EC and MBR method was
441 0.62 and showed a slope of 2, although the r^2 and slope decreased to 0.53 and 1.3,
442 respectively, when the regression was forced through the zero offset. The linear
443 regression for September 28 had an r^2 of 0.7 and a slope of 0.82 which decreased to 0.68
444 and 0.9, respectively, when forced through the zero offset. No significant linear
445 regression between EC and MBR measurements, however, was found on September 24
446 when most of the data fell into QC 3.

447 GEM fluxes measured with the MBR method showed an even higher range of values
448 and greater variability than to the EC measurements (Table 1). One possible reason for
449 this includes potentially different concentration footprints for the MBR method, which
450 has two different inlets at heights of 0.58 m and 1.09 m, respectively, compared to the EC
451 method (single inlet at 0.75 m). However, concentration footprint distances amounted to
452 8–10 m and 18–22 m for the two different gradient intake heights for typical afternoon
453 conditions and hence were clearly located within the 60 m source area. Therefore, the
454 different inlet heights of the EC and MBR methods cannot explain the observed
455 differences between the two methods or the more variable MBR flux data.

456

457 Conclusions

458 We successfully demonstrated applicability of the EC method for GEM flux
459 measurements by detection of well-defined covariance between high-resolution
460 measurements of vertical wind speed and GEM concentrations. Such field measurements
461 were challenging since they required deployment of a complex and highly sensitive
462 optical CRDS system in the field with significant demands for protecting the system from
463 vibrations, temperature and power fluctuations, and dust. Calculated GEM fluxes were in
464 the range of other flux methods under Hg enrichment, and showed high sensitivity to the
465 presence of Hg-enriched soil with changing fluxes as wind direction changed over the
466 source area. There was good comparison between the three flux measurement techniques
467 when conditions were favorable for flux measurements over the Hg-enriched source area
468 (QC 1 & 2). These measurements, to our knowledge, represent the first-ever EC
469 measurements performed to quantify GEM surface exchange. It proved to be
470 straightforward to detect EC fluxes over the Hg-enriched soil area, but detection was not
471 feasible over natural background areas in the existing setup with the current CRDS
472 sensitivity. The smallest, absolute, detectable flux with our system was $32 \text{ ng m}^{-2} \text{ h}^{-1}$, and
473 we conclude that this flux was a conservative estimate of the system's flux detection
474 limit. Various system improvements could be implemented to increase flux detection
475 limits and time coverage of measurements, including increasing the sensitivity of the
476 CRDS GEM detection limit, (e.g., by increasing cavity path lengths to improve
477 sensitivity), new cavity designs targeted at reducing flow interferences for EC
478 measurements (reduce noise evident in the GEM spectra and co-spectra and eliminate

479 reliance on a slow-response instrument for data correction), and advancements to the
480 system to make it more robust for longer-term field deployments.

481

482 **Figure Caption**

483 *Figure 1: a.) Site schematic; blue area represents the area with deployed Hg-enriched*
484 *soil boxes (represented by square dots), and dashed lines indicate the 10 m² spacing*
485 *around each individual source tub. b.) Flux measurement over background soils (left*
486 *panel) and over Hg-enriched areas (right panel) and associated flux footprints. Symbol*
487 *colors represent flux magnitude, and location show X₉₀% values, which represent upwind*
488 *distance from tower within which 90% of the flux originated, based on footprint analysis*
489 *(note micrometeorological tower located in the center). Solid symbols indicate QC 1*
490 *data, open symbols indicate QC 2 data, crossed out symbols indicate QC 3 data. The blue*
491 *area on the right panel indicates deployed Hg-enriched soil boxes. c.) Wind rose for the*
492 *measurement period.*

493

494 *Figure 2: a.) Power spectra of sonic temperature (T) and GEM (Hg) (left panel) and*
495 *corresponding co-spectra with vertical wind speed (right panel). Lines represent*
496 *ensemble averages from 19 September 2012 14-17:00 local time (mean wind speed 1.8*
497 *m s⁻¹ and unstable conditions). The dotted vertical line indicates the onset of noise (0.45*
498 *Hz). The solid line in the left panel shows the expected $f^{5/3}$ slope in the inertial subrange.*
499 *The solid and dashed lines in the right panel refer to the cospectral reference model by*
500 *Kaimal and Finnigan (1994) and the cospectral reference model attenuated by a series of*

501 *transfer functions that represent the effects of low-pass filtering on the GEM time series,*
502 *respectively.*

503

504 *Figure 3: Cross-correlation analysis of vertical wind speed with GEM density. The red,*
505 *blue and green lines refer to cases of a well-defined peak in the cross-correlation (QC*
506 *flag 1; 19.09.2012 14:00-14:30, average horizontal wind speed = 2.3 m s⁻¹, Monin-*
507 *Obukhov stability parameter = -0.15), a less well-defined peak (QC flag 2; 19.09.2012*
508 *17:30-18:00, average horizontal wind speed = 2.9 m s⁻¹, Monin-Obukhov stability*
509 *parameter = -0.04) and no detectable peak (QC flag 3; 20.09.2012 16:30-17:00, average*
510 *horizontal wind speed = 2.9 m s⁻¹, Monin-Obukhov stability parameter = -0.13).*

511

512 *Figure 4: GEM fluxes (ng m⁻² h⁻¹) as measured by the CRDS system (filled dots) and the*
513 *MBR method (open dots) during three specific days with different wind directions and*
514 *therefore different impact of the deployed Hg-enriched sources. Shaded areas represent*
515 *the angles corresponding to the deployed Hg-enriched substrate upwind of the*
516 *measurement towers; wind directions are average wind directions during the respective*
517 *30 min. measurement periods and show the degree to which dominant wind directions*
518 *corresponded to the upwind Hg-enriched source area. Wind direction error bars are*
519 *standard deviation of the wind direction.*

520

521

522 Supporting Information

523 The supporting information includes further information of the field site and
524 meteorology, data processing, auxiliary instruments, flux footprints and flux
525 comparisons. This material is available free of charge via the Internet at
526 <http://pubs.acs.org>.

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532 Author Contributions

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547

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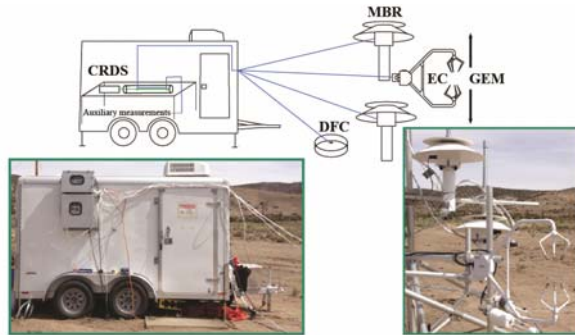
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